

Number of Neutrino Types

The neutrinos referred to in this section are those of the Standard $SU(2) \times U(1)$ Electroweak Model possibly extended to allow nonzero neutrino masses. Light neutrinos are those with $m < m_Z/2$. The limits are on the number of neutrino mass eigenstates, including ν_1 , ν_2 , and ν_3 .

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Number from $e^+ e^-$ Colliders

Number of Light ν Types

VALUE	DOCUMENT ID	TECN
2.9840 ± 0.0082	¹ LEP-SLC	06 RVUE
• • • We do not use the following data for averages, fits, limits, etc. • • •		
3.00 ± 0.05	² LEP	92 RVUE

¹ Combined fit from ALEPH, DELPHI, L3 and OPAL Experiments.

² Simultaneous fits to all measured cross section data from all four LEP experiments.

Number of Light ν Types from Direct Measurement of Invisible Z Width

In the following, the invisible Z width is obtained from studies of single-photon events from the reaction $e^+ e^- \rightarrow \nu \bar{\nu} \gamma$. All are obtained from LEP runs in the E_{cm}^{ee} range 88–209 GeV.

VALUE	DOCUMENT ID	TECN	COMMENT
2.92 ± 0.05 OUR AVERAGE	Error includes scale factor of 1.2.		
2.84 ± 0.10 ± 0.14	ABDALLAH	05B DLPH	$\sqrt{s} = 180\text{--}209$ GeV
2.98 ± 0.05 ± 0.04	ACHARD	04E L3	1990–2000 LEP runs
2.86 ± 0.09	HEISTER	03C ALEP	$\sqrt{s} = 189\text{--}209$ GeV
2.69 ± 0.13 ± 0.11	ABBIENDI,G	00D OPAL	1998 LEP run
2.89 ± 0.32 ± 0.19	ABREU	97J DLPH	1993–1994 LEP runs
3.23 ± 0.16 ± 0.10	AKERS	95C OPAL	1990–1992 LEP runs
2.68 ± 0.20 ± 0.20	BUSKULIC	93L ALEP	1990–1991 LEP runs
• • • We do not use the following data for averages, fits, limits, etc. • • •			
2.84 ± 0.15 ± 0.14	ABREU	00Z DLPH	1997–1998 LEP runs
3.01 ± 0.08	ACCIARRI	99R L3	1991–1998 LEP runs
3.1 ± 0.6 ± 0.1	ADAM	96C DLPH	$\sqrt{s} = 130, 136$ GeV

Limits from Astrophysics and Cosmology

Effective Number of Light ν Types

(“Light” means $<$ about 1 MeV). The quoted values correspond to N_{eff} , where $N_{\text{eff}} = 3.046$ in the Standard Model with $N_\nu = 3$. See also OLIVE 81. For a review of limits based on Nucleosynthesis, Supernovae, and also on terrestrial experiments, see DENEGRI 90. Also see “Big-Bang Nucleosynthesis” in this *Review*.

<u>VALUE</u>	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
● ● ● We do not use the following data for averages, fits, limits, etc. ● ● ●				
3.3 ± 0.5	95	¹ ADE	14	COSM Planck
$3.78^{+0.31}_{-0.30}$		² COSTANZI	14	COSM
3.29 ± 0.31		³ HOU	14	COSM
< 3.80	95	⁴ LEISTEDT	14	COSM
< 4.10	95	⁵ MORESCO	12	COSM
< 5.79	95	⁶ XIA	12	COSM
< 4.08	95	MANGANO	11	COSM BBN
$0.9 < N_\nu < 8.2$		⁷ ICHIKAWA	07	COSM
$3 < N_\nu < 7$	95	⁸ CIRELLI	06	COSM
$2.7 < N_\nu < 4.6$	95	⁹ HANNESTAD	06	COSM
$3.6 < N_\nu < 7.4$	95	⁸ SELJAK	06	COSM
< 4.4		¹⁰ CYBURT	05	COSM
< 3.3		¹¹ BARGER	03C	COSM
$1.4 < N_\nu < 6.8$		¹² CROTTY	03	COSM
$1.9 < N_\nu < 6.6$		¹² PIERPAOLI	03	COSM
$2 < N_\nu < 4$		LISI	99	COSM BBN
< 4.3		OLIVE	99	COSM BBN
< 4.9		COPI	97	Cosmology
< 3.6		HATA	97B	High D/H quasar abs.
< 4.0		OLIVE	97	BBN; high ^4He and ^7Li
< 4.7		CARDALL	96B	COSM High D/H quasar abs.
< 3.9		FIELDS	96	COSM BBN; high ^4He and ^7Li
< 4.5		KERNAN	96	COSM High D/H quasar abs.
< 3.6		OLIVE	95	BBN; ≥ 3 massless ν
< 3.3		WALKER	91	Cosmology
< 3.4		OLIVE	90	Cosmology
< 4		YANG	84	Cosmology
< 4		YANG	79	Cosmology
< 7		STEIGMAN	77	Cosmology
		PEEBLES	71	Cosmology
< 16		¹³ SHVARTSMAN	69	Cosmology
		HOYLE	64	Cosmology

¹ Fit to the number of neutrino degrees of freedom from Planck CMB data along with WMAP polarization, high L, and BAO data.

² Fit to the number of neutrinos degrees of freedom from Planck CMB data along with BAO, shear and cluster data.

³ Fit based on the SPT-SZ survey combined with CMB, BAO, and H_0 data.

⁴ Constrains the number of neutrino degrees of freedom (marginalizing over the total mass) from CMB, CMB lensing, BAO, and galaxy clustering data.

⁵ Limit on the number of light neutrino types from observational Hubble parameter data with seven-year WMAP data, SPT, and the most recent estimate of H_0 . Best fit is 3.45 ± 0.65 .

⁶ Limit on the number of light neutrino types from the CFHTLS combined with seven-year WMAP data and a prior on the Hubble parameter. Best fit is $4.17^{+1.62}_{-1.26}$. Limit is relaxed to $3.98^{+2.02}_{-1.20}$ when small scales affected by non-linearities are removed.

⁷ Constrains the number of neutrino types from recent CMB and large scale structure data. No priors on other cosmological parameters are used.

AKERS	95C	ZPHY C65 47	R. Akers <i>et al.</i>	(OPAL Collab.)
OLIVE	95	PL B354 357	K.A. Olive, G. Steigman	(MINN, OSU)
BUSKULIC	93L	PL B313 520	D. Buskulic <i>et al.</i>	(ALEPH Collab.)
LEP	92	PL B276 247	LEP Collabs.	(LEP, ALEPH, DELPHI, L3, OPAL)
WALKER	91	APJ 376 51	T.P. Walker <i>et al.</i>	(HSCA, OSU, CHIC+)
DENEGRI	90	RMP 62 1	D. Denegri, B. Sadoulet, M. Spiro	(CERN, UCB+)
OLIVE	90	PL B236 454	K.A. Olive <i>et al.</i>	(MINN, CHIC, OSU+)
YANG	84	APJ 281 493	J. Yang <i>et al.</i>	(CHIC, BART)
OLIVE	81	APJ 246 557	K.A. Olive <i>et al.</i>	(CHIC, BART)
OLIVE	81C	NP B180 497	K.A. Olive, D.N. Schramm, G. Steigman	(EFI+)
STEIGMAN	79	PRL 43 239	G. Steigman, K.A. Olive, D.N. Schramm	(BART+)
YANG	79	APJ 227 697	J. Yang <i>et al.</i>	(CHIC, YALE, UVA)
STEIGMAN	77	PL 66B 202	G. Steigman, D.N. Schramm, J.E. Gunn	(YALE, CHIC+)
PEEBLES	71	Physical Cosmology	P.Z. Peebles	(PRIN)
		Princeton Univ. Press (1971)		
SHVARTSMAN	69	JETPL 9 184	V.F. Shvartsman	(MOSU)
		Translated from ZETFP 9 315.		
HOYLE	64	NAT 203 1108	F. Hoyle, R.J. Tayler	(CAMB)
